

First steps towards cosmological simulations with full EAGLE physics

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with

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Cosmological scales?



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Any needs for more?

- Thinking of future weak-lensing surveys:
 - Measure some cosmological information on scales down to
 ~ 1-30 Mpc. Clearly "baryon effects" seen on these scales.
 - "Common wisdom" asks for volume in excess of 300 Mpc.
 - That asks for particles counts in excess of 4500³ ~ 100 billion.
 - With EAGLE code that would be >300M CPU hours and 1.3PB of RAM

SWIFT 101

- Aimed at replacing Gadget for EAGLE-like runs.
- Use task-based parallelism, modern algorithms, better parallelisation and domain decomposition.
- Leaner memory footprint, faster i/o, more modular, multiple hydro schemes.
- Collaboration with computer scientists and industry.

SWIFT 101

- Hydro neighbour finding based on regular AMR cell structure. Many flavours of SPH + "GIZMO".
- FMM for gravity with a multipole-mesh method for periodic gravity.
- Particles sorted to enhance the vectorization of the code.
- Activation of work only in the "active" parts of the tree.

SWIFT interlude - AGORA galaxy



L. Hausammann, Y. Revaz, MS



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EAGLE code performance



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Strong scaling behaviour



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AMR Verlet-list neighbour search



- Target ~500 particles per cell via adaptive mesh refinement.
- Cell size naturally matches particle neighbour search radius.
- Particles only interact with particles in the same cell or any direct neighbouring cell.



- Cells pairs do not need to be processed in any pre-defined order.
- Only need to make sure two threads do not work on the same cell.
- Cell pairs can have vastly different work-loads.



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→ Need runtime dynamic scheduling

Task-based parallelism for science case



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Task-based parallelism in action



Task-graph for one time-step. Colours correspond to different task types. Almost perfect load-balance achieved on 16 cores.

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Local time-stepping

Our particle methods for gravity and hydro-dynamics have a linear cost.

 \rightarrow By only updating the particles on a short time-step we can gain orders of magnitude in run time.

 \rightarrow You also kill your scaling.

The question is then how to load-balance and parallelize this. How do you update ~10 particles efficiently on 1000+ nodes?

Domain Decomposition

- 1. Sort the data along a spacefilling curve.
- 2. Split the curve such that each node gets the same number of particles.
- 3. Do this at the top-level or close to the top.



Smallest time-steps?



A hand-written solution

- Avoid MPI for small updates.
- The particles with the smallest time-steps should be at the centre of their domains and not require any "halo" particles.
- Small time-steps are at the centre of galaxies.
- Identify galaxies → Grow domains organically around them
 → Get efficient code.

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Drawbacks: Not generic. Identifying galaxies is not trivial.

A Graph-based strategy



- For each task, we compute the amount of work (=runtime) required.
- We build a graph where the data are nodes and tasks are hyper-edges.
- METIS is used to split the graph such that the work (not the data!) is balanced.
- Extra cost added for communication tasks to minimise them.

What does it look like?



- No regular grid pattern.
- No space-filling curve pattern.
- Good (work) load-balancing by construction.
- The most dense regions are at the centre of their respective domains.

SWIFT code performance



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Strong scaling behaviour



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Strong scaling behaviour



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Sustained performance in weak-scaling



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Conclusions

• New algorithms can lead to significant speed-ups over more conventional established methods.

>30x over Gadget → "half way from peta-scale to exa-scale via algorithms"

- Task-based parallelism as a viable model for actual scientific applications.
 - \rightarrow Not just a research concept.
- The challenge of deep time-step hierarchies requires a rethinking of the domain decomposition algorithms.

